

Nuclear data uncertainty propagation: Total Monte Carlo vs. covariances

D. Rochman, A.J. Koning, S.C. van der Marck,
and D. van Veen

Nuclear Research and Consultancy Group,

NRG, Petten, The Netherlands

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Contents



① Goals:

\implies *Propagate nuclear data uncertainties for the SG-33*

② Methodology for uncertainty propagation:

\implies *TMC vs. covariances (exact or with covariances ?)*

③ Models:

\implies *(1) Total Monte Carlo and (2) perturbation*

④ Tests:

\implies *Consistence between both methods*

⑤ Preliminary results:

\implies *on k_{eff} for $^{239,240}\text{Pu}$, pmf1 and pmf2*

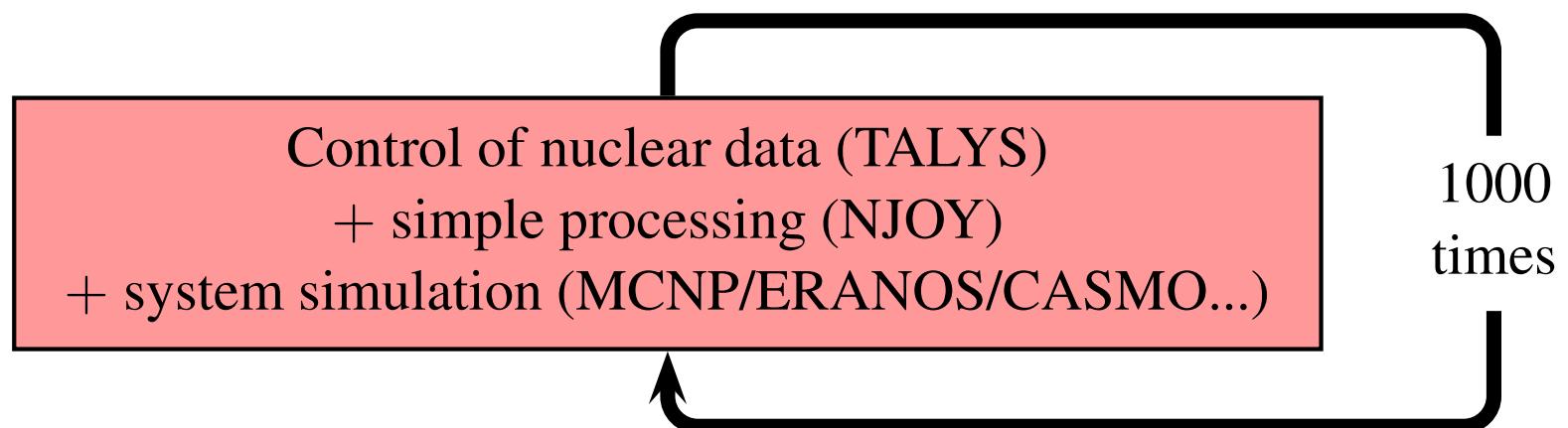
⑥ Conclusions

Goals:



-
- ① Obtain uncertainties for SG-33 due to nuclear data uncertainties
 - ② Systematic approach, reliable and reproducible

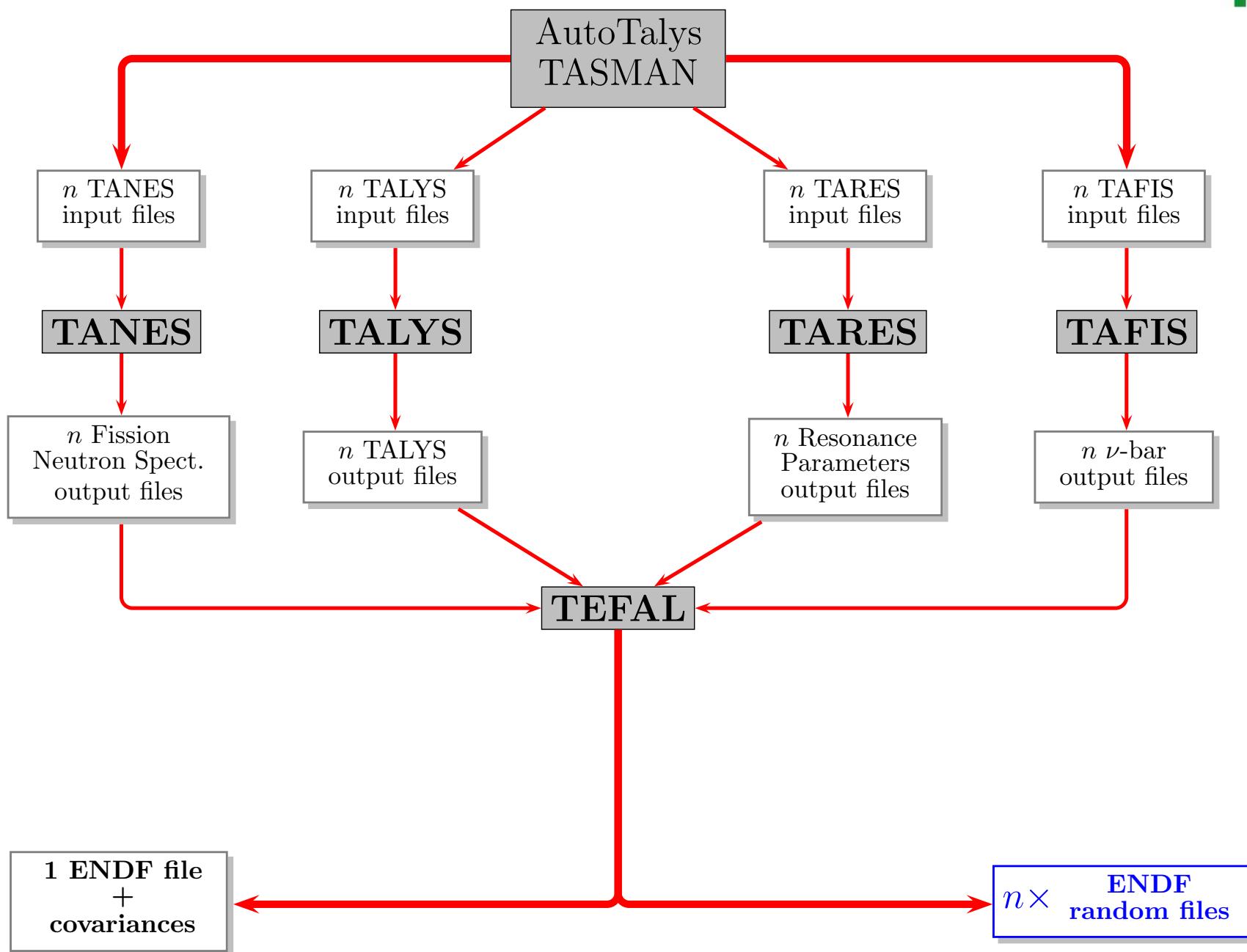
Solution (1): Total Monte Carlo



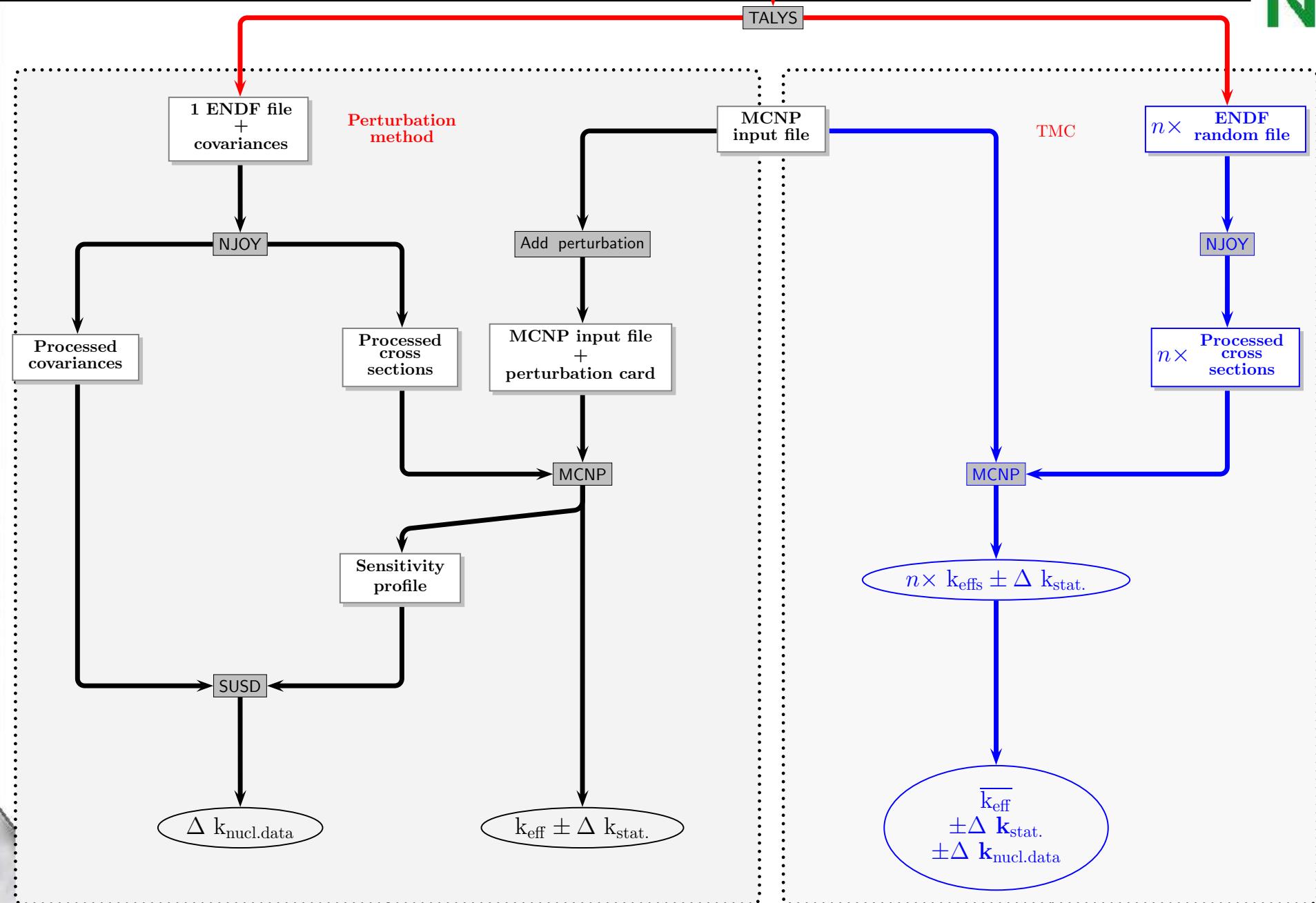
Solution (2): Perturbation method

⇒ MCNP + Perturbation cards + covariance files

TMC and Perturbation method: File production



TMC and Perturbation method



Necessary softwares



Common to TMC and Perturbation methods:

- ☞ TALYS
- ☞ NJOY (ACE)
- ☞ MCNP

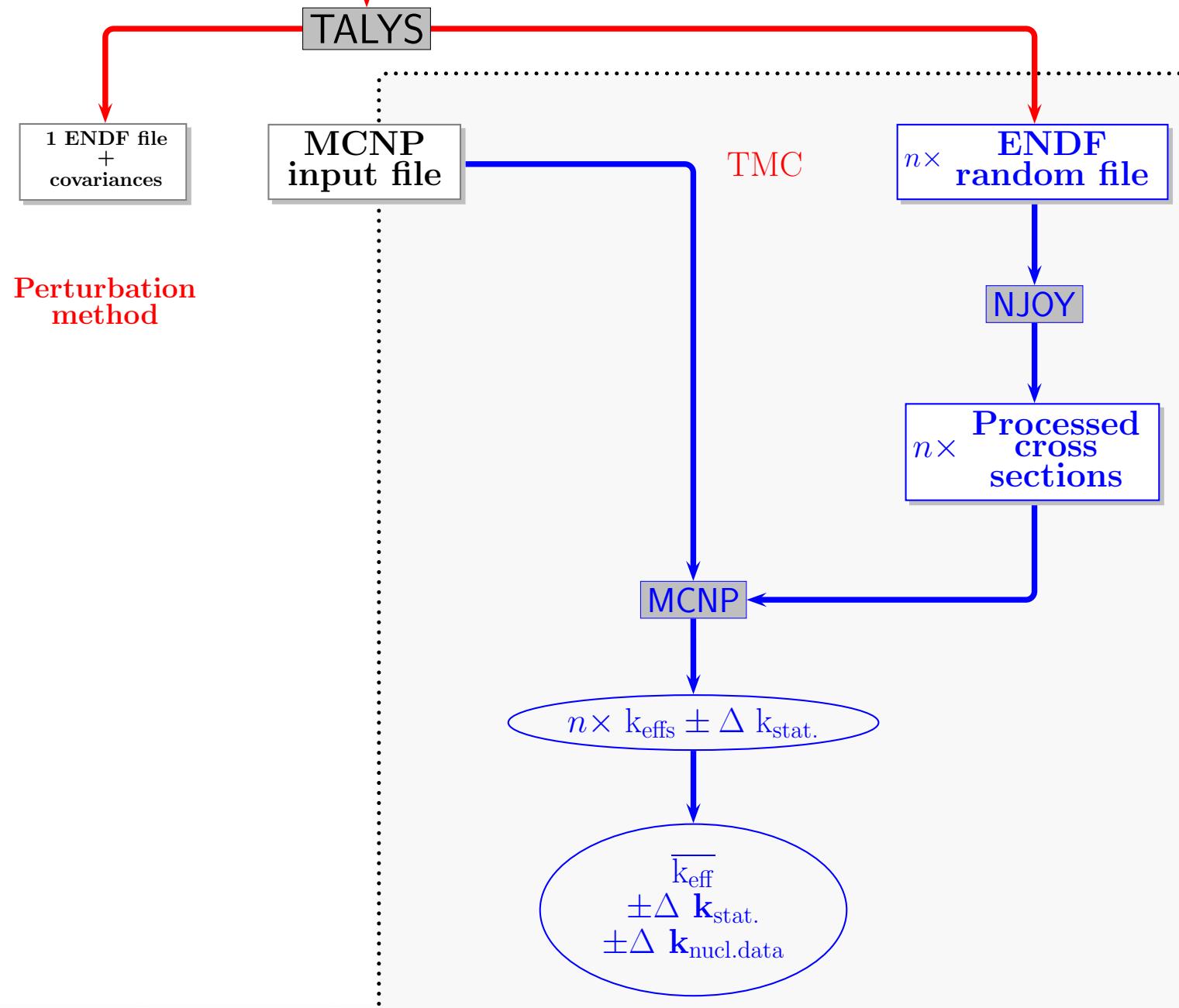
For the TMC method only:

∅

For the Perturbation method only:

- ☞ NJOY (ERRORR)/PUFF
- ☞ Add perturbation
- ☞ SUSD

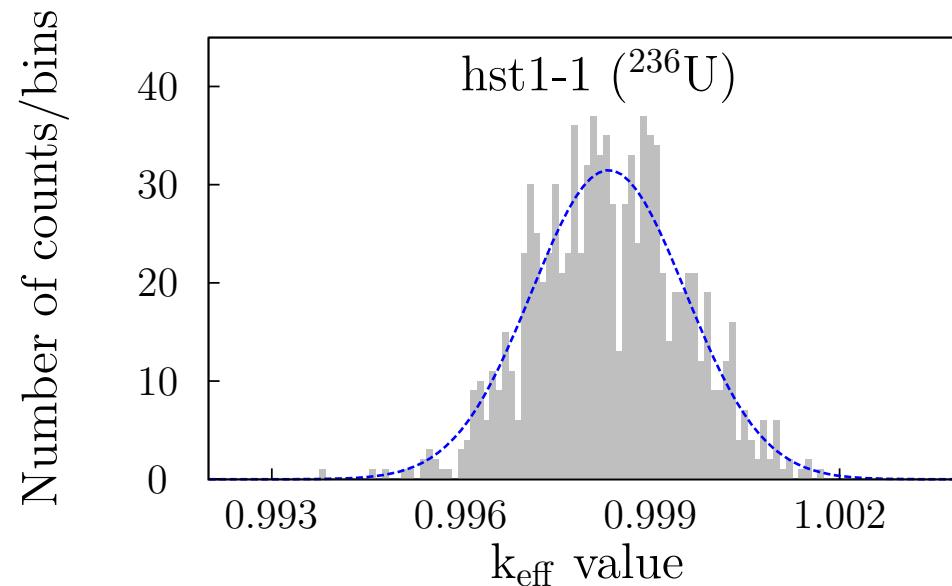
Idea: TALYS +(n TALYS input files) Monte Carlo = Total Monte Carlo



Total Monte Carlo: examples

For each random ENDF file, the benchmark calculation is performed with MCNP. At the end of the n calculations, n different k_{eff} values are obtained. In the obtained probability distribution of k_{eff} , the standard deviation σ_{total} reflects two different effects:

$$\sigma_{\text{total}}^2 = \sigma_{\text{statistics}}^2 + \sigma_{\text{nuclear data}}^2. \quad (1)$$



Each random file is completely different than another one: nu-bar ("MF1"), resonance parameters ("MF2"), cross sections ("MF3"), but also *MF4*, *MF5* and *MF6*.

Perturbation method



n TALYS
input files

TALYS

1 ENDF file
+ covariances

Perturbation
method

MCNP
input file

NJOY

Processed
covariances

Processed
cross
sections

MCNP input file
+ perturbation card

Add perturbation

MCNP

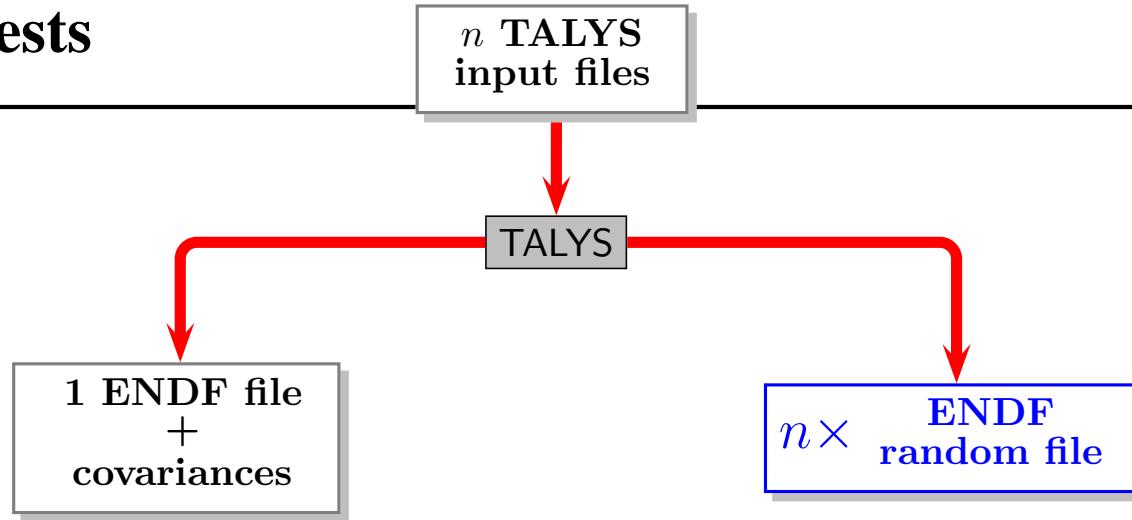
Sensitivity
profile

SUSD

$\Delta k_{\text{nucl.data}}$

$k_{\text{eff}} \pm \Delta k_{\text{stat.}}$

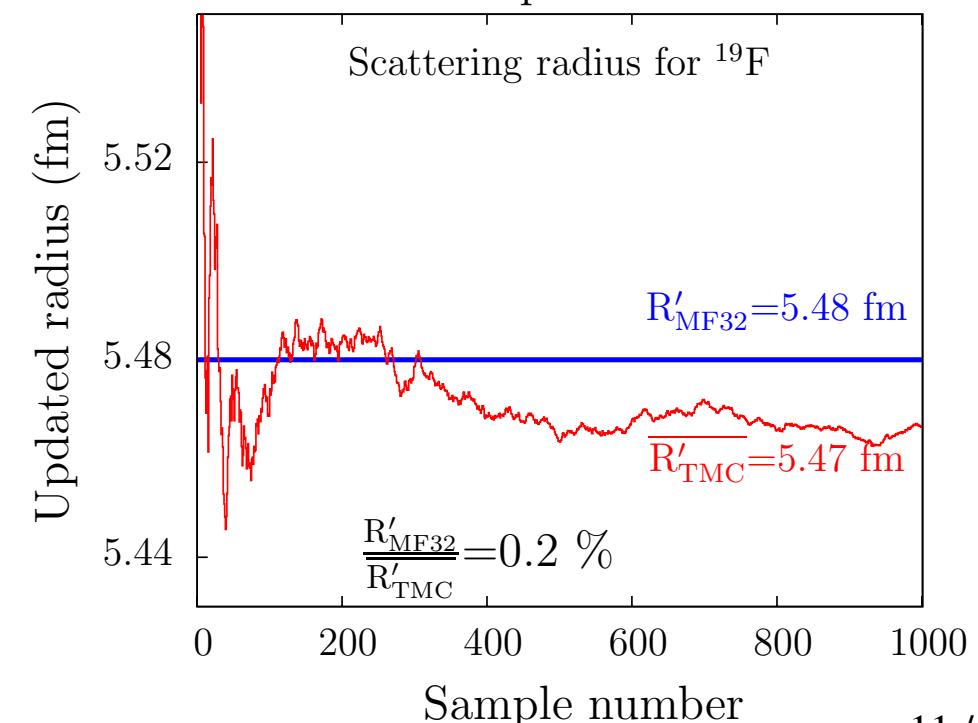
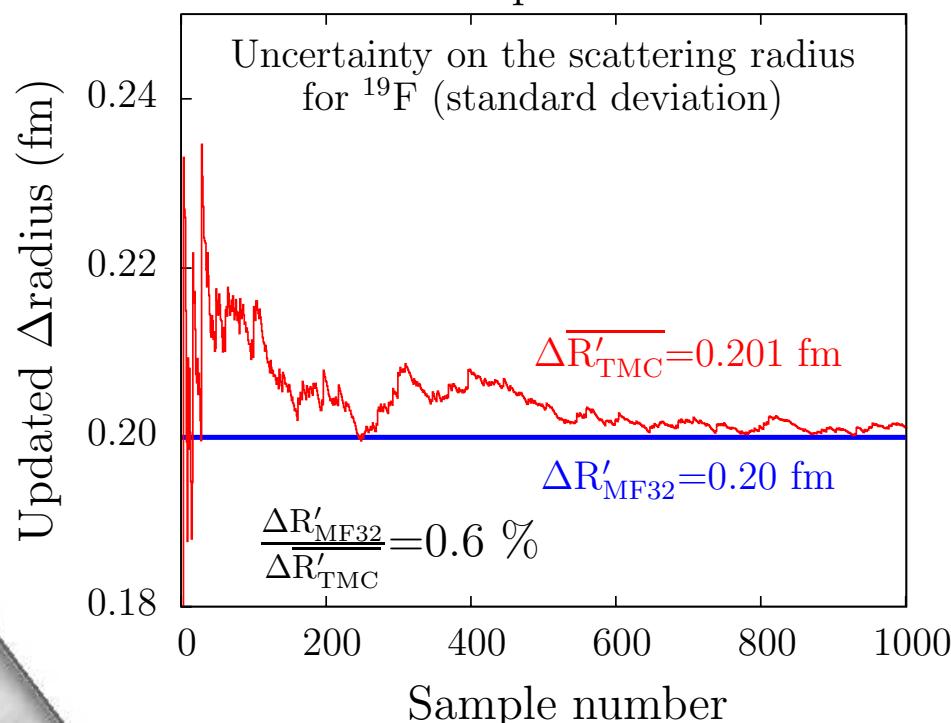
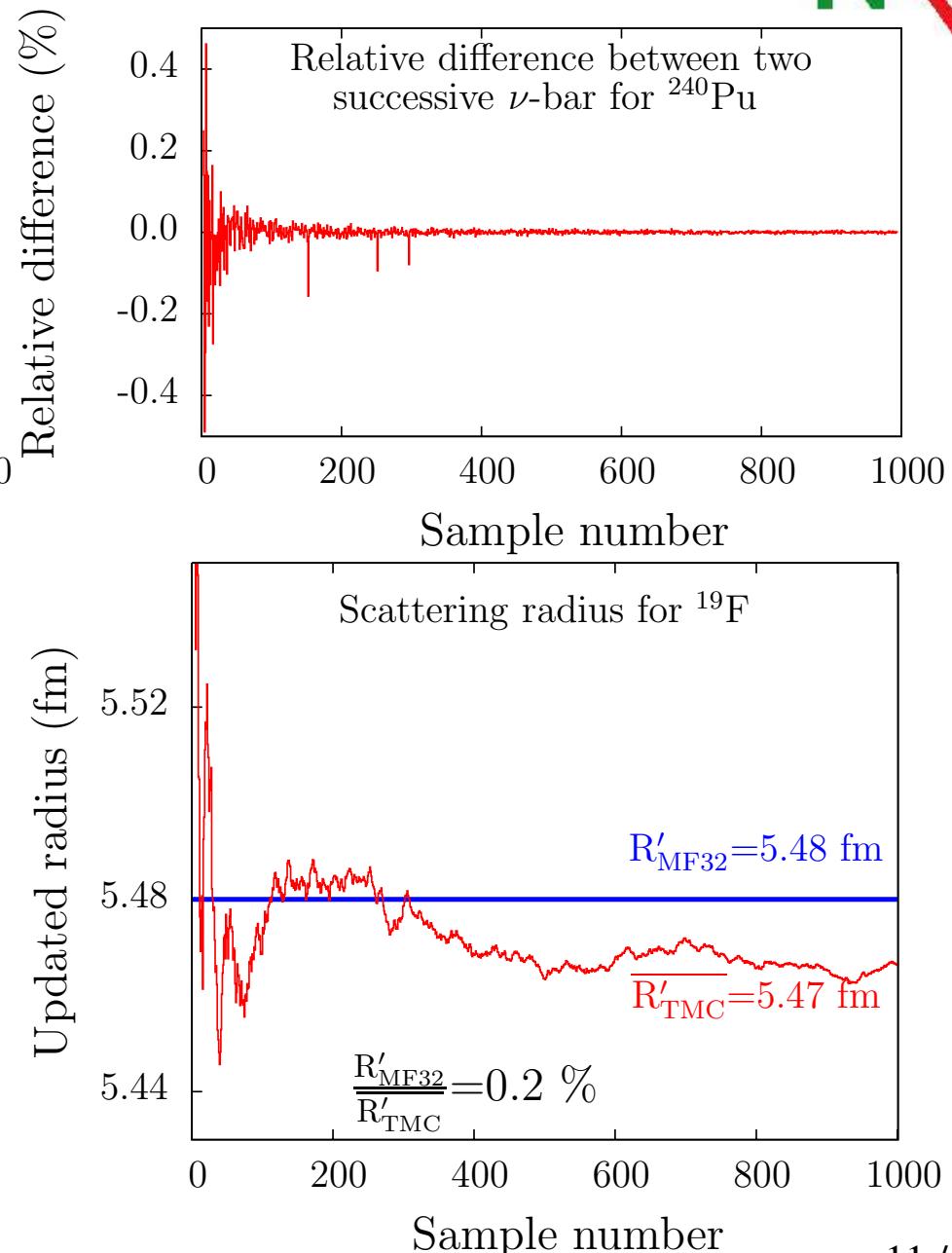
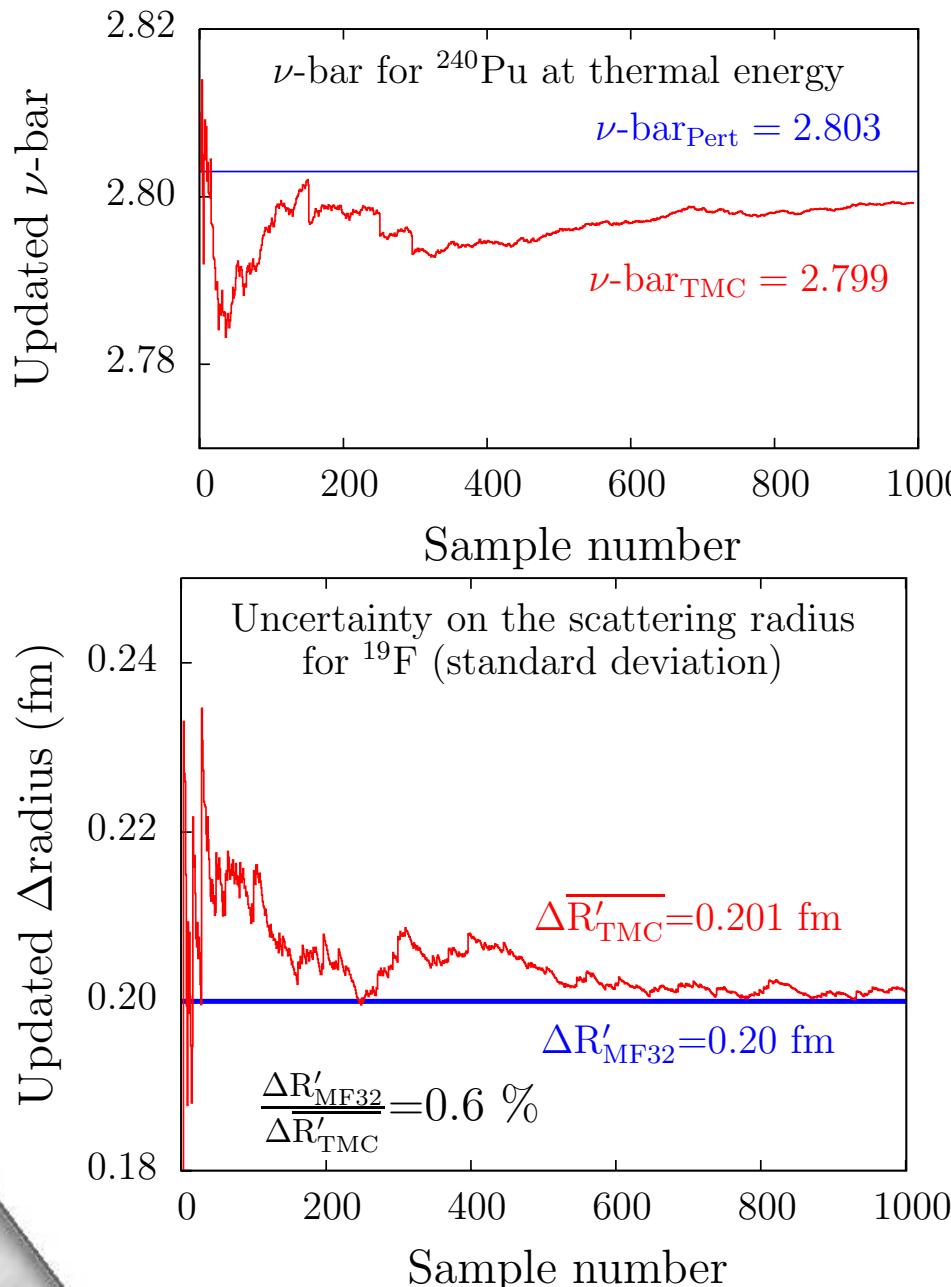
Consistency tests



Convergence and consistency of:

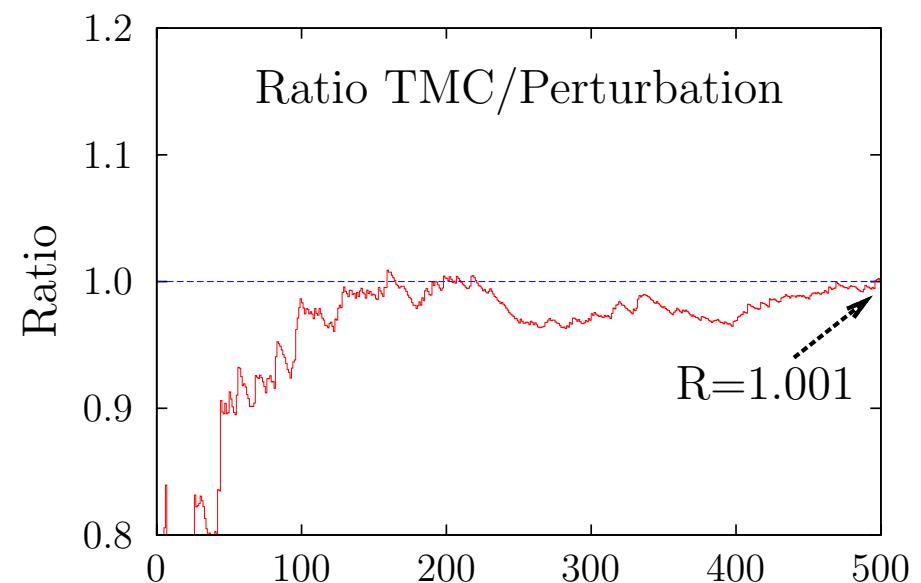
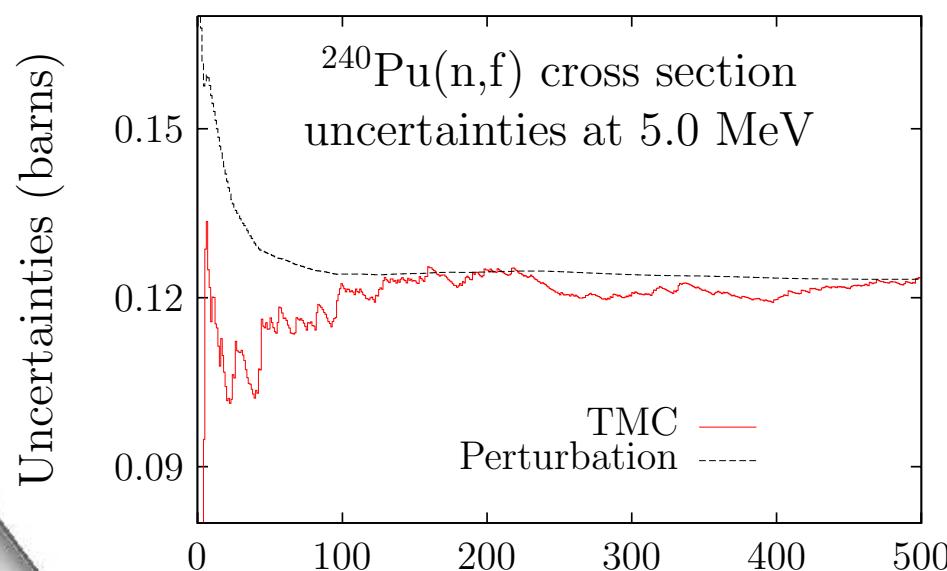
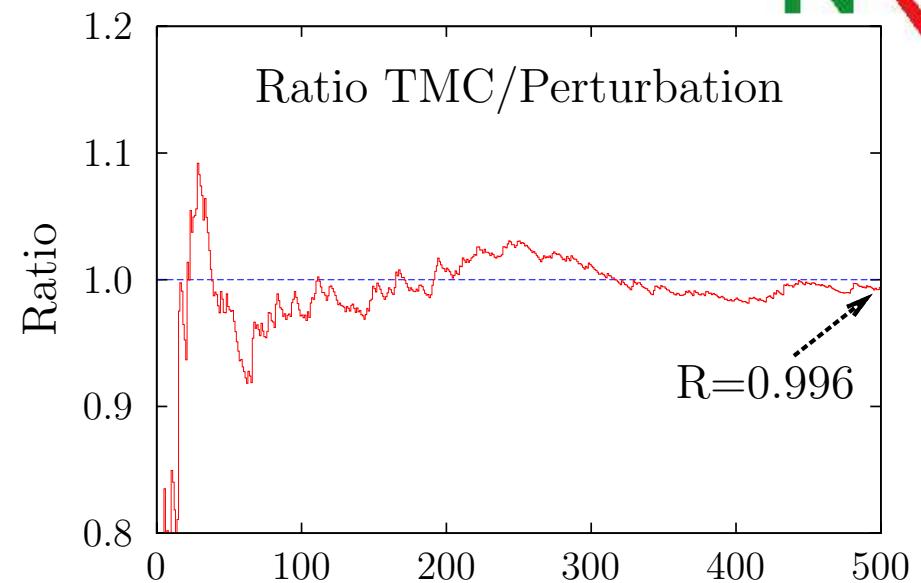
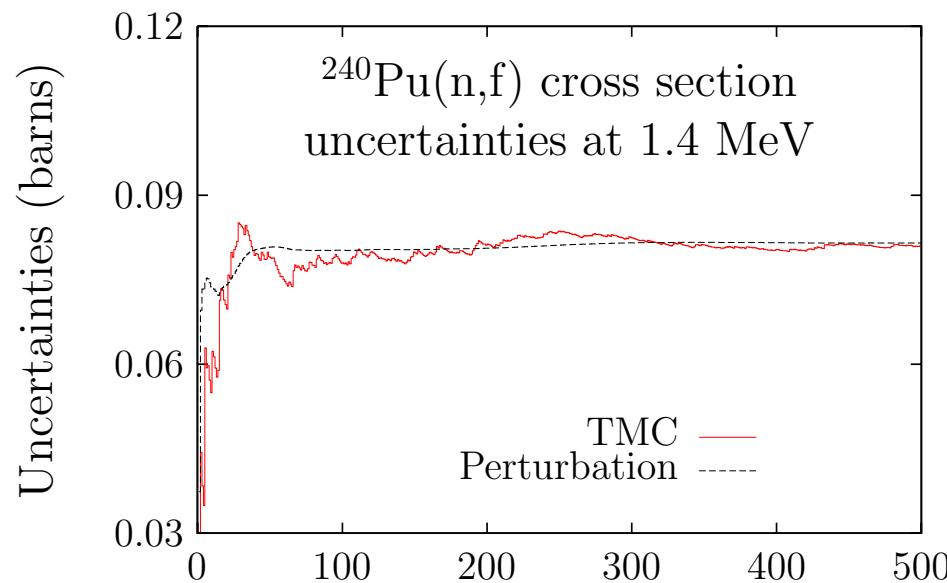
- ❖ $\bar{\nu}$.
- ❖ resonance parameter distributions.
- ❖ cross sections probability distributions.
- ❖ angular distribution probability distributions.
- ❖ Monte Carlo calculations.
- ❖ the perturbation method.

Convergence and consistency of ν -bar and resonance parameters

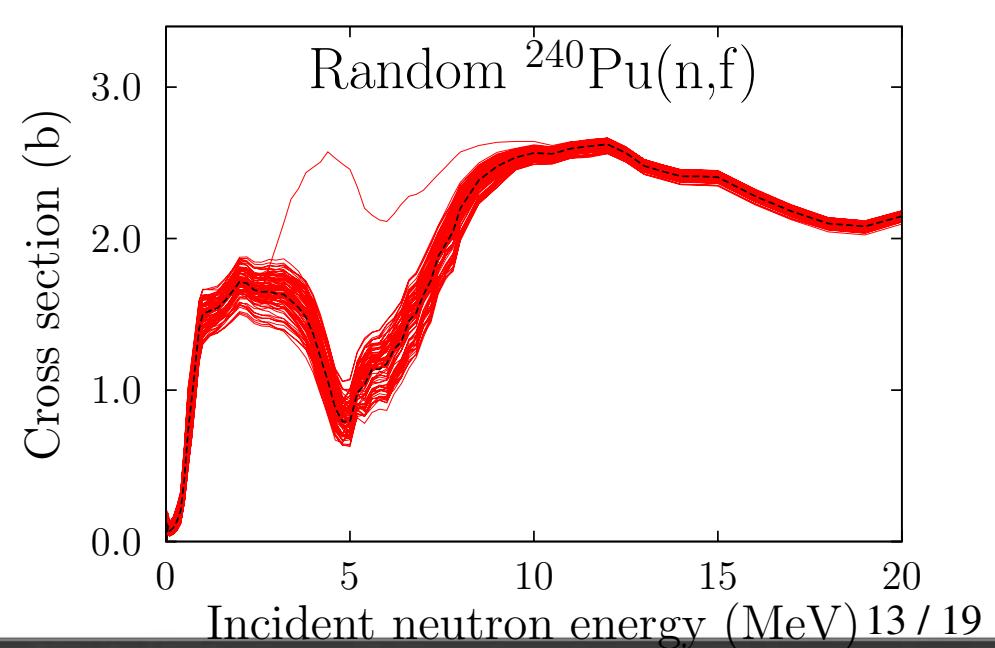
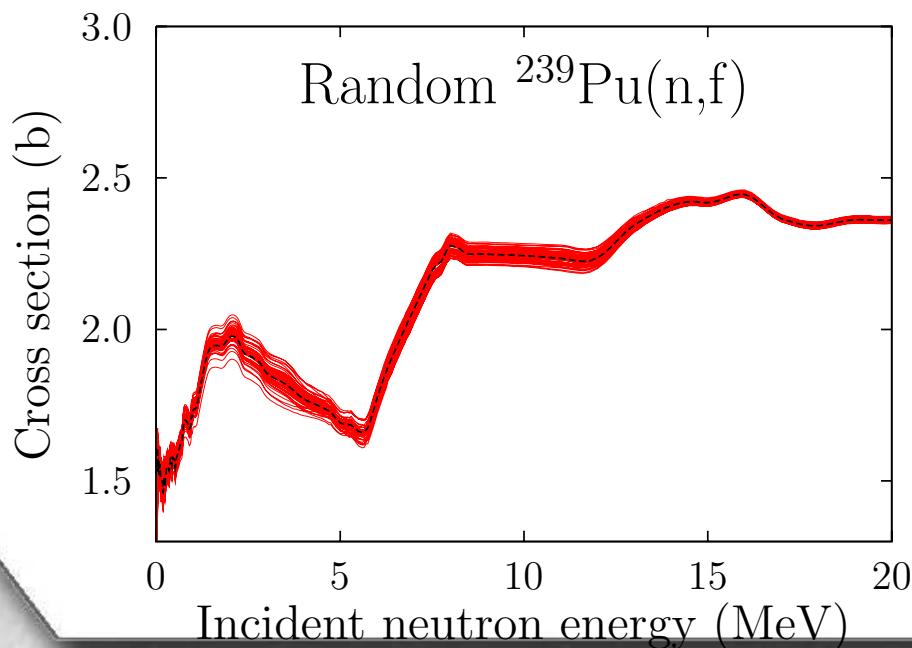
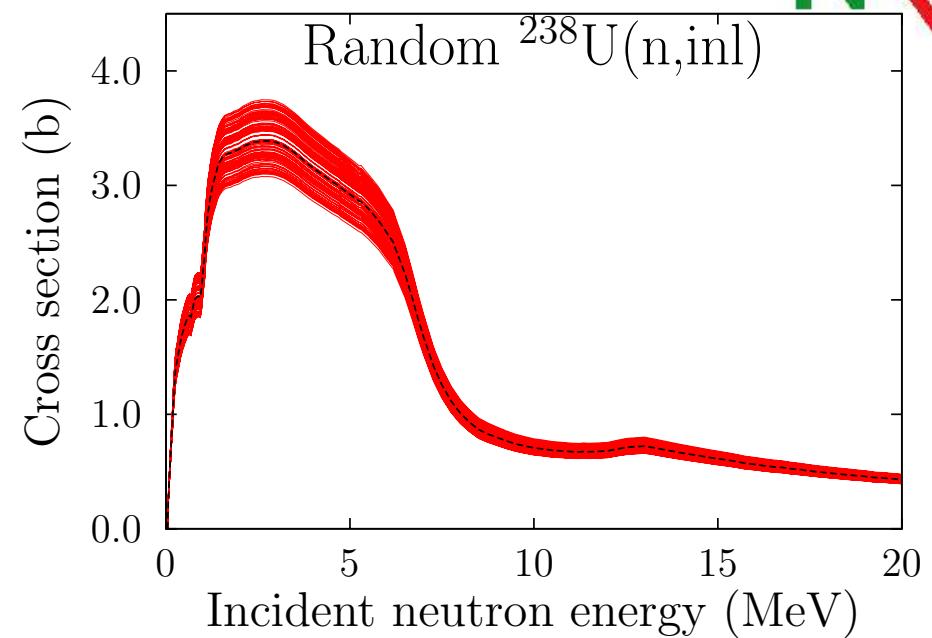
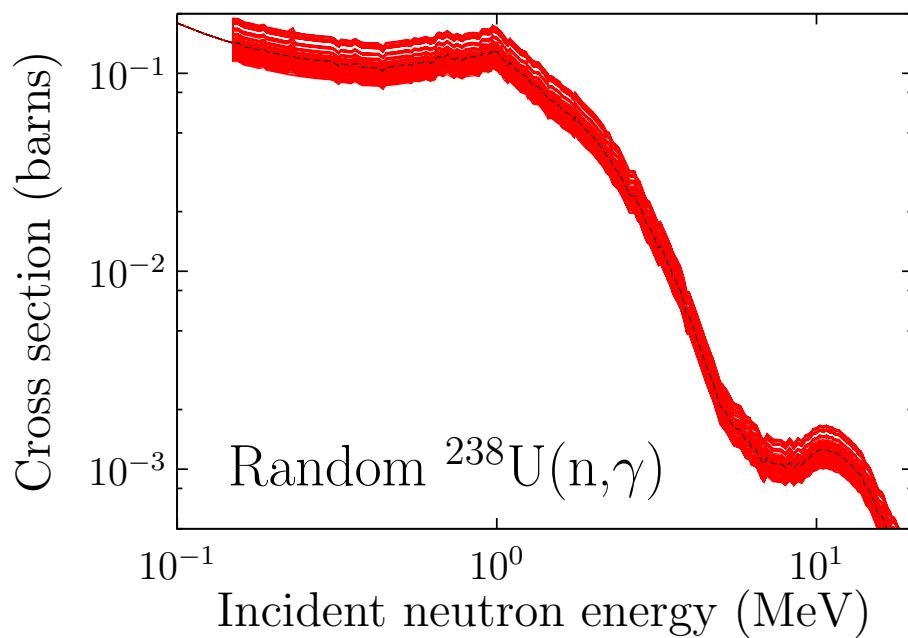


Convergence TMC/Perturbation

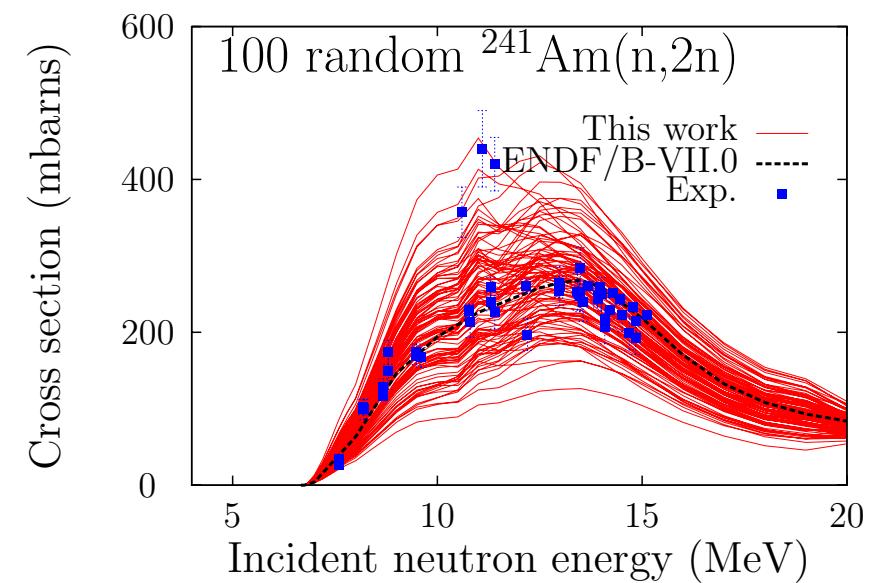
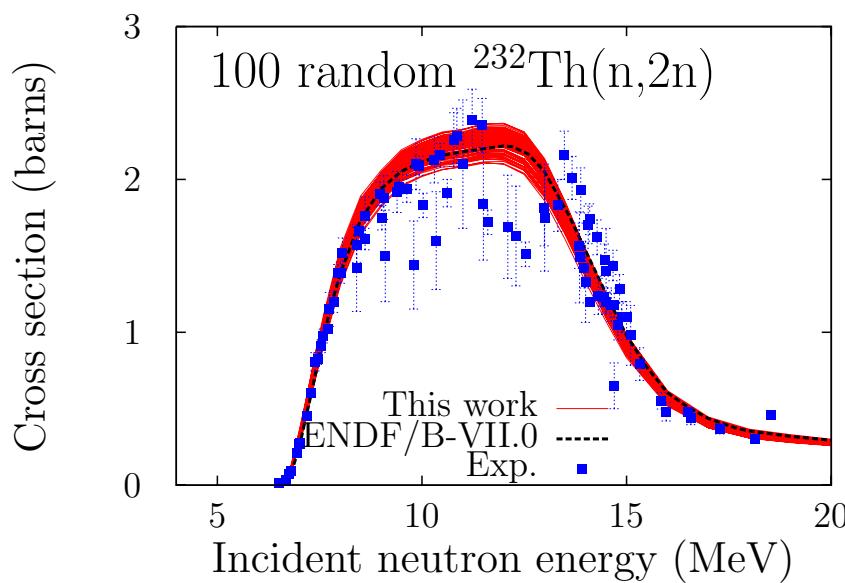
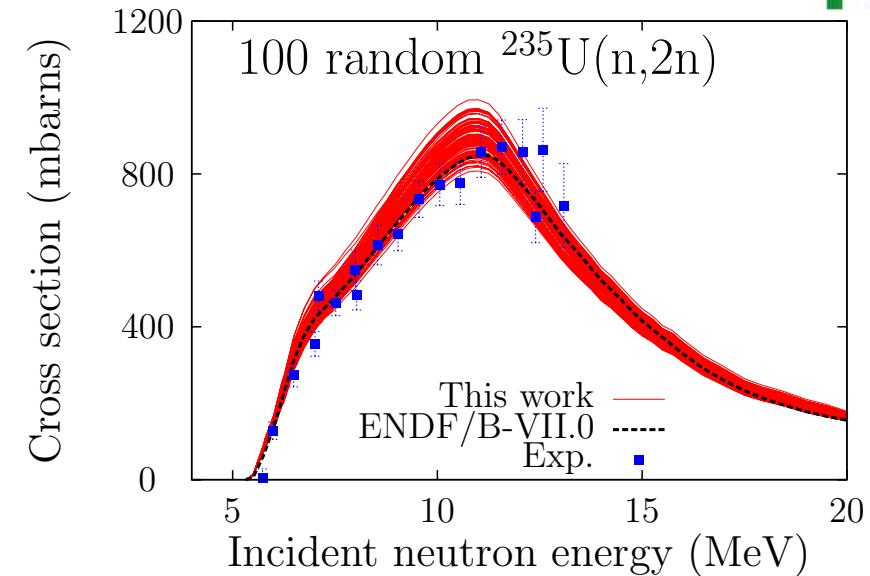
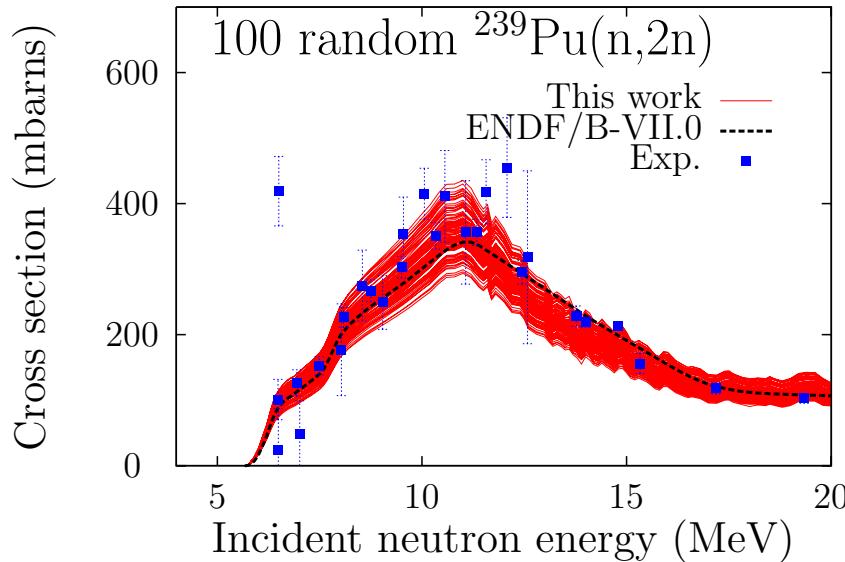
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Nuclear data: ^{239}Pu and ^{238}U

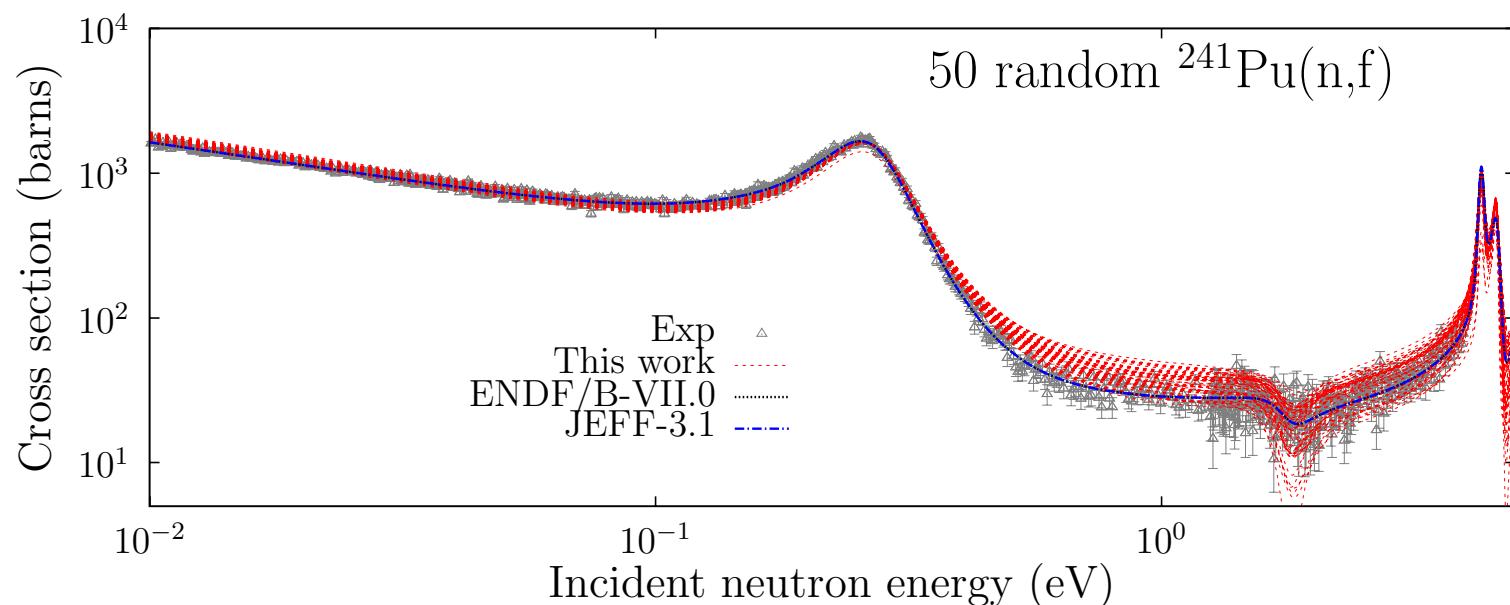
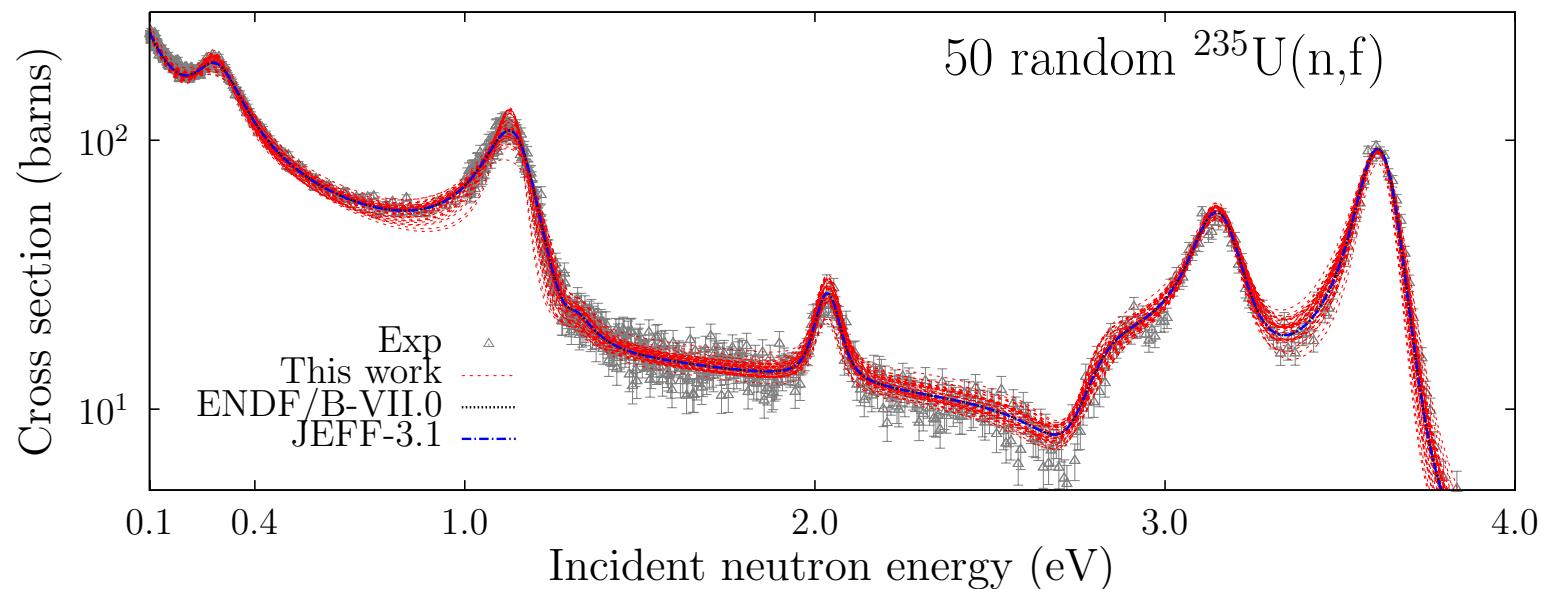


Nuclear data: examples on (n,2n) cross sections



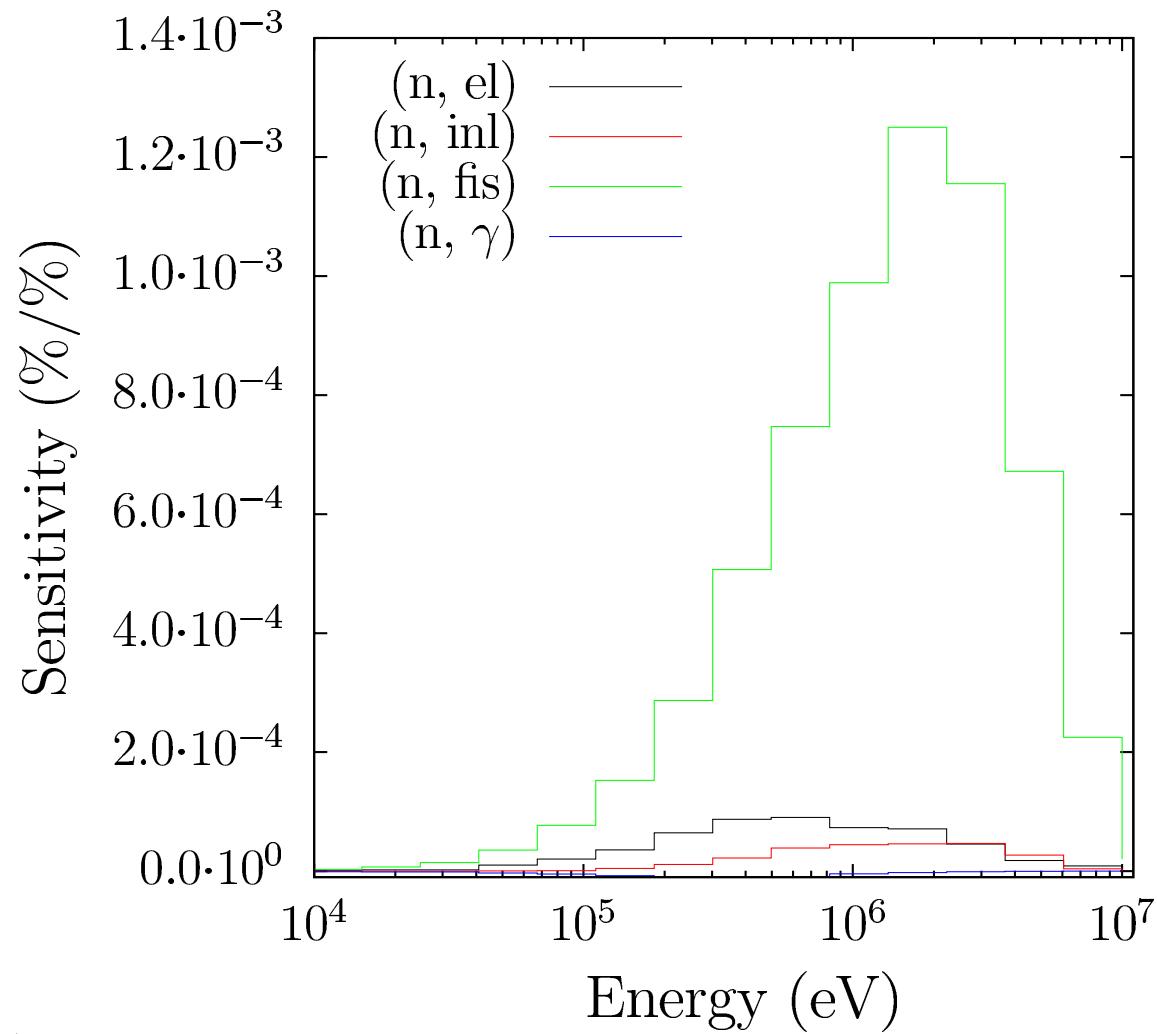
Nuclear data: examples in the resonance region

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Example for sensitivity to ^{239}Pu cross section for pmf1 (Jezebel)

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Results



Comparison TMC-Perturbation methods for a few k_{eff} benchmarks. The ratio in the last column is "TMC over Perturbation".

Benchmark	Isotopes	Total Monte Carlo Uncertainty due to nuclear data (pcm)	Perturbation Uncertainty due to nuclear data (pcm)	Ratio
pmf1	^{239}Pu	1000	860	1.16
pmf2	^{239}Pu	840	720	1.16
pmf2	^{240}Pu	790	650	1.21

Results: Details of the TMC-Perturbation methods for $^{239,240}\text{Pu}$ k_{eff} benchmarks



	pmf2 ^{239}Pu			pmf2 ^{240}Pu	
	Δk_{eff} (pcm)			Δk_{eff} (pcm)	
	TMC	Perturbation		TMC	Perturbation
Total	840	720		790	650
MF1	400	-		370	-
(n,inl)	170	140		70	50
(n,el)	250	240		30	40
(n, γ)	100	100		30	30
(n,f)	720	660		730	640
MF4	20	-		20	-
MF5	50	-		30	-
MF6	50	-		30	-

Conclusions



- 😊 First attempt to compare two uncertainty propagation method
- 😊 TMC: more general and exact answer, does not require special codes, more exhaustive
- 🙁 but slower

- 🙁 Perturbation: approximate, require special processing and codes, limited
- 😊 but faster
- ⌚ TMC uncertainties \simeq 15-20 % larger than from perturbation for pmf2 considering $^{239,240}\text{Pu}$